

Linewidth Analyzer

High resolution lineshape spectra and ultra sensitive noise analyzers for narrowband lasers

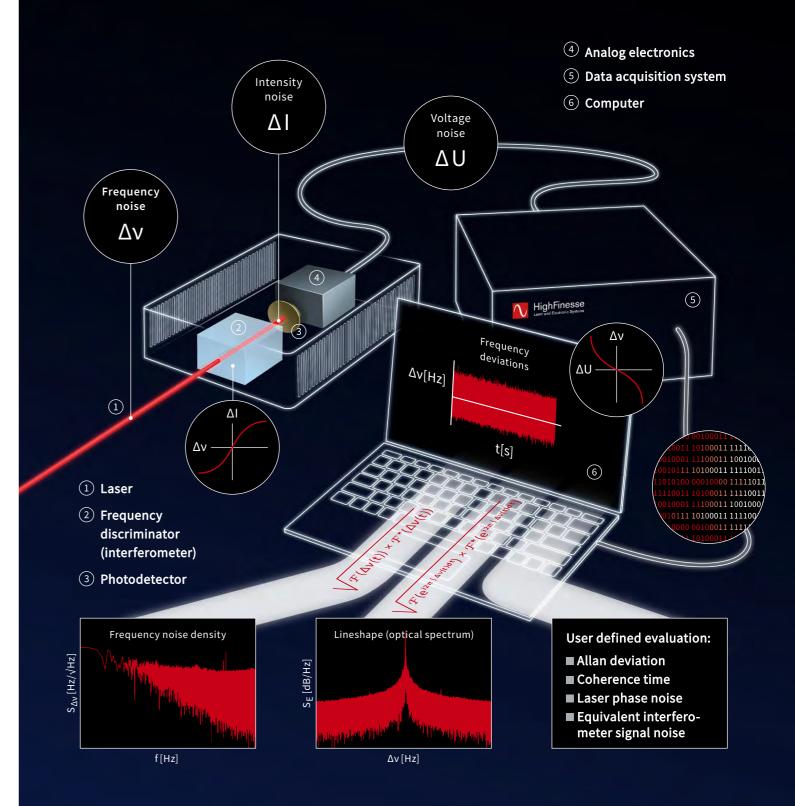
LWA Series

The HighFinesse Linewidth Analyzers are the ultimate high-end instruments for measuring, analyzing and controlling frequency and intensity noise of lasers. The superb sensitivity of these instruments is achieved by combining an interferometric working principle with high-end optical and electronic components.

The main features are:

- Frequency noise density, optical lineshape and relative intensity noise (RIN) analysis with evaluation of intrinsic (Lorentzian) and effective (optical) linewidth
- Intrinsic (Lorentzian) linewidth measurements down to 350 Hz
- Effective (optical) linewidth measurement range down to 1 kHz
- Frequency noise density floor down to 5 Hz/√Hz with a dynamic range of 60 dB between 10 Hz and 10 MHz
- Laser phase noise density floor down to 3.2 µrad/√Hz
- RIN measurement down to -150 dB/Hz
- Robust against acoustic noise
- Error signal generator for further linewidth, frequency noise or RIN reduction
- Powerful tool for a detailed analysis of noise sources like servo bumps, frequency drifts, power supply noise and acoustics
- Extremely fast analysis: up to real time measurements and evaluation





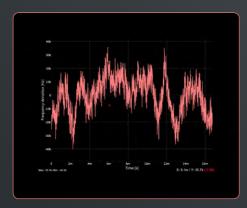
The laser light ① is coupled to the input fiber and lead through an interferometer ② acting as a frequency discriminator. The transmitted intensity, which is directly proportional to the variations of the input frequency, is converted by a photodetector ③ and our analog electronics ④ into a voltage signal. This voltage is finally digitized by the Digitizer ⑤ to provide the data for evaluation on a computer ⑥.

The included software recovers the original frequency noise using the precisely known interferometer function.

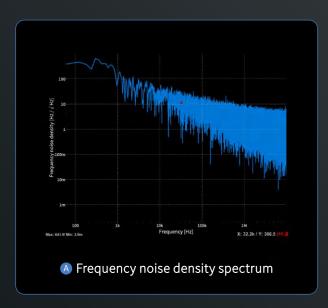
The recovered timeseries of the frequency deviations is now the basic dataset allowing to calculate easily the frequency noise density spectrum and the optical lineshape spectrum.

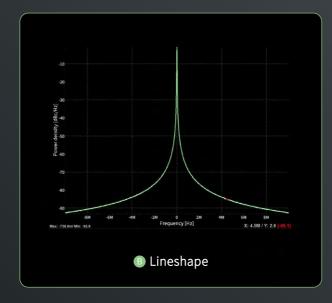
The user can also export the timeseries data in order to perform custom evaluation methods such as Allan deviation or coherence time analysis. Let the Software Suite do the math for you. The LWA Software Suite is delivered with all LWA products.

Method (A)
Fourier analysis of the
frequency deviations
Frequency noise



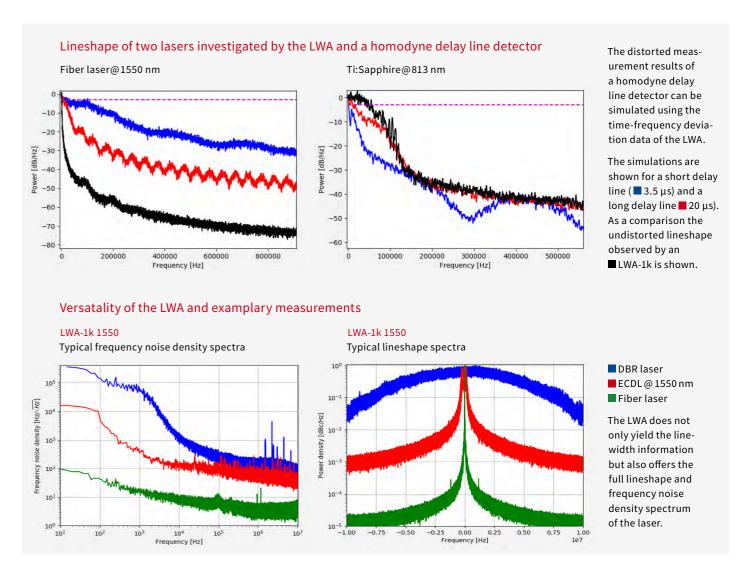
Method B
Calculation based on
the Fourier analysis
of the electrical field
Lineshape





- Real-Time Analysis
- Automatic calculation of:
- frequency noise density.
- intrinsic and effective laser linewidth based on the β-separation approach [DiDomenico Appl. Opt. 49 (2010)].
- optical lineshape for various observation times.
- linewidth parameters by fitting Voigt, Lorentzian or Gaussian models to the observed lineshape.
- Timeseries export allows for calculating many mode parameters like Allan deviation, coherence time, equivalent interferometer signal noise, etc.

Beside the frequency discrimination approach, homo- and heterodyne delay-line techniques are commonly used for linewidth determination. However, using optical delay lines can lead to complex spectra with non-trivial evaluation needs due to the technique inherent loss of information.



Discriminator (e.g. LWA-1k)

Principle
Limits
Basic data

Frequency noise density
Lineshape

Linewidth evaluation

Laser phase noise spectrum

Interferometer signal noise spectrum

Direct frequency noise to intensity conversion

Steepness of discriminator-function

Frequency deviations in time

By a Fourier analysis of the frequency deviations

By a Fourier analysis of the calculated electric fields

Effective (optical) and intrinsic (Lorentzian) linewidths

directly accessible via frequenvy density noise spectrum or lineshape spectrum

By performing a Fourier analysis of the integrated frequency fluctuations

By performing a Fourier analysis of the calculated interferometer phase signal ¹⁾

Delay-line approach

Optical delayed superposition by different path lengths

Length of delay path

Spectrum of beat note (convolution of original spectrum)

Not possible without additional methods

Lorentzian part (intrinsic linewidth) of the lineshape

Only the Lorentzian part of the linewidth is accessible, because the frequency noise is high-passed by the delay line

Not possible without additional methods

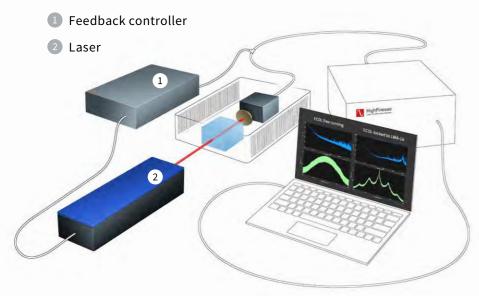
Not possible without additional methods

Assuming the interferometer working purely as a frequency discriminator and the interferometer phase signal being proportional to the frequency deviation signal.

Feedback Controller

Due to the design of the LWAs, the output voltage can be directly used as an error signal for a feedback controller allowing to reduce the frequency noise of the test laser.

Depending on the used feedback controller and the laser system the optical linewidth can be reduced by more than two orders of magnitude offering a vast amount of new opportunities.



Active laser noise reduction

- Connect the Analyzer output signal (a) as input signal to a fast feedback controller.
- Connect the feedback controller to the laser's fast DC modulation input (e.g. laser diode current).

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- Adjust the feedback to minimize the output signal of the Analyzer (e.g. PID parameters, gain)

Typical application

- Laser module quality control
- Laser design optimization
- Metrology and quantum technology
- Linewidth control for spectroscopy
- Modulation surveillance

Product Overview

Wavelength range	
nput power range (@typical wavelengt	h)
Required input power stability	
_aser type	
nput fiber type	
Maximum frequency stroke (@ f > 10Hz)	
Frequency Noise Specification	
Noise floor $N_{\Delta v}$ @ typ. input power and wavelength $^{8)}$	
aser phase noise floor @typ. input pound wavelength 4) 9)	ver
Equivalent interferometer signal noise @typ. input power and wavelength ^{4) 6)}	
requency noise bandwidth 5)	
Minimum measurable intrinsic linewidt lorentzian linewidth)	h
Effective linewidth range (optical linewi β-separation method]	dth)
Relative intensity noise limit lorentzian linewidth)	
Dynamic range	
ineshape Specifications	
Effective linewidth range (optical linewi curve fitting method]	idth)
Dynamic range	
Miscellaneous	
nterface	
Analog Output	
Cutoff (highpass filter)	
Dimensions	
Veight	
Digitizer Module	
Sample rate	
Resolution	
Acquisition time (time series)	
Evaluation time 7)	
Communication	

²⁾ Frequency noise and lineshape specifications are derived from measurements at 780 nm.

3) PM fiber recommended.

Weight

- 4) Not included in the software, can be calculated by the user from exported data.
- 5) According to a -3 dB criterion.

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Unit		LW			LWA-10k VIS ²⁾						LWA-1k 1550							/A-1	0k 1	NIR		LWA-100k NIR											
	min		typ max		ax	min		ty	typ max		ax	min		typ		max		min		typ		max		m	nin typ		max						
nm			78	30	10	64	45	50	7	80	10	64	15	30	15	50	16	525	10	54	15	50	16	525	10	64	15	50	16	525			
mW			10		1	15		0.5		5		8		.5	5		8		0.	5	į	5		8	0.	.5	5		8				
%	_														±	5																	
													Lase	r typ	e CV	I, sin	gle r	node															
		F	PM-F	C/AP	C		SM or PM ³⁾ , FC/AP						PM-FC/APC							SM or PM ³⁾ , FC/AP						SM or PM ³⁾ , FC/AP							
MHz	30						40						30							40							100						
————	10	100	1k	10k	100k	>1M	10	100	1k	10k	100k	>1M	10	100	1k	10k	100k	>1M	10	100	1k	10k	100k	>1M	10	100	1k	10k	100k	>1M			
		_				_		_		_				_				_						_		_				_			
Hz/√Hz	200	75	30	30	25	15	500	150	60	60	50	30	80	40	15	10	8	5	200	100	30	20	15	10	1k	200	60	50	40	25			
rad/√Hz	20	750 m	30 m	3 m	250 μ	15μ	5	1.5	60 m	6 m	500μ	30 μ	8	400 m	15 m	1 m	80μ	5μ	20	1	30 m	2 m	150 μ	10 μ	100	2	60 m	5 m	400 μ	25 μ			
dBrad/√Hz	26	-2.5	-30	-50	-72	-96	24	3.5	-24	-44	-66	-90	18	-8	-36	-60	-82	-106	26	0	-30	-54	-76	-100	40	6	-24	-46	-68	-92			
rad/√Hz/m	6.2 μ	2.3μ	920 n	920 n	770 n	460 n	16 μ	4.6 μ	1.8μ	1.8 μ	1.6μ	920 n	2.5 μ	1.3μ	460 n	310 n	250 n	160 n	6.2 μ	3.1 μ	920 n	620 n	460 n	310 n	31 μ	7μ	2μ	2μ	2μ	765 n			
dBrad/√Hz/m	-104	-112	-120	-120	-122	-126	-96	-106	-114	-114	-116	-120	-112	-118	-126	-130	-132	-136	-104	-110	-120	-124	-126	-130	-90	-104	-115	-116	-118	-122			
Hz															10 -	10 M																	
Hz	<3 k							<12 k						<350						<2 k						<10 k							
Hz		<	10 k	- 20	М		< 20 k - 30 M							<1 k - 20 M							< 5 k - 30 M						<15 k - 100 M						
dB/Hz	-														-150							-						-					
dB												60																					
Hz	<10 k - 10 M							< 20 k - 10 M							< 1 k - 10 M							< 5 k - 10 M						< 15 k - 10 M					
dB												60																					
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mm	220 × 334 × 96 440 × 34									5 mn	—— n	220 × 334 × 96						440 × 340 × 155 mm						440 × 340 × 155 mm									
kg				8.0 12						8.0							12							12									
Sa/s														62	.5 M	(ma	x.)																
bits															1	6																	
s														1	m –	100	n																
S														10	m –	1 (ty	p.)																
															B 3.0																		
mm														21	0 × 2		74																
kg															2	.0																	

⁶⁾ This is the calculated noise of the interferometer phase of a two path interferometer with length imbalance L (in meters). The alculation is performed for a given frequency noise density floor by $2\pi n L/c \times N_{\Delta v}$ with n being the refractive index of the reference fiber interferometer material and c being the speed of light in vacuum. Values in the table are given for an refractive index of n = 1.46 and a reference length of 1 meter.

 $Linewidth Analyzer \cdot 2-2024 \cdot This \ document \ provides \ general \ information only \ and \ may \ be \ subject to \ change \ at \ any time \ without \ prior \ notice.$

⁷⁾ Windows 10 or newer, Intel i5 8600/AMD Ryzen 5 2600 or better, 16GB RAM or more.

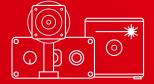
N_{Av} is the noise floor of the instrument in terms of the square root of the power spectral density of the frequency noise.

⁹⁾ The phase noise floor corresponds to the noise floor of the square root of the power spectral density of the phase. It is calculated from N_{Δν} by the formula 1/f × N_{Δν}. Additionally, phase noise is often specified in terms of L(f) which can be calculated with the formula L(f) = 1/f² × N²_{Δν}/2.









Wavelength Meter

HighFinesse/Ångstrom offers sensitive and compact wavelength meters with a large spectral range for high speed measurement of lasers. The optical unit consists of temperature-controlled Fizeaubased interferometers that are read out by photodiode arrays. The high absolute accuracy is achieved by use of solid state, non-moving optics. The optical unit and associated electronics are housed in a compact, thermal casing. The connection to a computer or notebook is realized via a highspeed USB 2.0 port, which allows a high data read-out rate. The analyzing software displays all the interferometer information.

Spectrometer

The grating based HighFinesse/ Ångstrom Laser Spectrum Analyzers offer the capability for a very accurate simultaneous measurement of both the center wavelength and the linewidth of a laser source with a compact and versatile instrument.

The product series covers the ranges from 192 nm to 2250 nm. The grating based technology allows the analysis of laser sources over a large linewidth range. Utilizing the principle of non-moving parts just like the well-known HighFinesse WS-series wavemeters, the LSA offers the time-tested robustness and ability to measure both pulsed and cw lasers.

Calibration Sources

HighFinesse offers a variety of frequency stabilized, narrow linewidth laser sources for the calibration of all wavelength meters and applications down to \pm 0.5 MHz absolute accuracy. These are user friendly, plug and play devices that can be connected to the wavelength meter. Different technologies, accuracies and wavelengths are available to suit your application.

HighFinesse stabilized frequency references yield extremely accurate frequency stabilizations, ideal for calibration of our wavelength meters in the visible and infrared wavelength regimes.



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